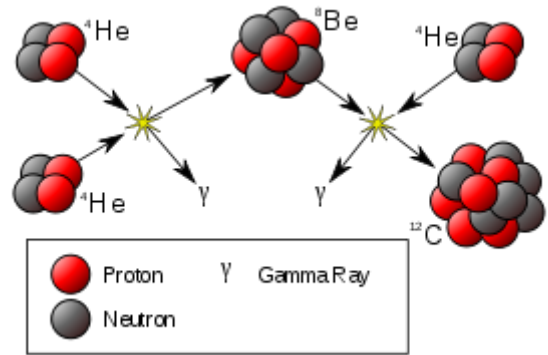


Triple-alpha process

The **triple-alpha process** is a set of nuclear fusion reactions by which three helium-4 nuclei (alpha particles) are transformed into carbon.^{[1][2]}

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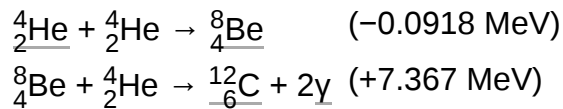
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Overview of the triple-alpha process.

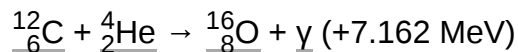
Triple-alpha process in stars

Helium accumulates in the core of stars as a result of the proton–proton chain reaction and the carbon–nitrogen–oxygen cycle. Further nuclear fusion reactions of helium with hydrogen or another alpha particle produce lithium-5 and beryllium-8 respectively. Both products are highly unstable and decay, almost instantly, back into smaller nuclei, unless a third alpha particle fuses with a beryllium-8 nucleus before that time to produce a stable carbon-12 nucleus. The half-life of ⁵Li is 3.7×10^{-22} s and that of ⁸Be is 8.19×10^{-17} s.^[3] When a star runs out of hydrogen to fuse in its core, it begins to contract and heat up. If the central temperature rises to 10^8 K,^[4] six times hotter than the Sun's core, alpha particles can fuse fast enough to produce significant amounts of carbon:



The net energy release of the process is 7.275 MeV.

As a side effect of the process, some carbon nuclei fuse with additional helium to produce a stable isotope of oxygen and energy:



Fusing with additional helium nuclei can create heavier elements in a chain of stellar nucleosynthesis known as the alpha process, but these reactions are only significant at higher temperatures and pressures than in cores undergoing the triple-alpha process. This creates a situation in which stellar nucleosynthesis produces large amounts of carbon and oxygen but only a small fraction of those elements are converted into neon and heavier elements. Oxygen and carbon make up the main "ash" of helium-4 burning.

Primordial carbon

Because the triple-alpha process is unlikely, it normally needs a long time to produce much carbon. One consequence of this is that no significant amount of carbon was produced in the Big Bang because, within minutes after the Big Bang, the temperature fell below the critical point for nuclear fusion.

Resonances

Ordinarily, the probability of the triple alpha process is extremely small. However, the beryllium-8 ground state has almost exactly the energy of two alpha particles. In the second step, ${}^8\text{Be} + {}^4\text{He}$ has almost exactly the energy of an excited state of ${}^{12}\text{C}$. This resonance greatly increases the probability that an incoming alpha particle will combine with beryllium-8 to form carbon. The existence of this resonance was predicted by Fred Hoyle before its actual observation, based on the physical necessity for it to exist, in order for carbon to be formed in stars. The prediction and then discovery of this energy resonance and process gave very significant support to Hoyle's hypothesis of stellar nucleosynthesis, which posited that all chemical elements had originally been formed from hydrogen, the true primordial substance. The anthropic principle has been cited to explain the fact that nuclear resonances are sensitively arranged to create large amounts of carbon and oxygen in the universe.^{[5][6]}

Nucleosynthesis of heavy elements

With further increases of temperature and density, fusion processes produce nuclides only up to nickel-56 (which decays later to iron); heavier elements (those beyond Ni) are created mainly by neutron capture. The slow capture of neutrons, the s-process, produces about half of elements beyond iron. The other half are produced by rapid neutron capture, the r-process, which probably occurs in core-collapse supernovae and neutron star mergers.^[7]

Reaction rate and stellar evolution

The triple-alpha steps are strongly dependent on the temperature and density of the stellar material. The power released by the reaction is approximately proportional to the temperature to the 40th power, and the density squared.^[8] In contrast, the proton–proton chain reaction produces energy at a rate proportional to the fourth power of temperature, the CNO cycle at about the 17th power of the temperature, and both are linearly proportional to the density. This strong temperature dependence has consequences for the late stage of stellar evolution, the red giant stage.

For lower mass stars on the red giant branch, the helium accumulating in the core is prevented from further collapse only by electron degeneracy pressure. The entire degenerate core is at the same temperature and pressure, so when its mass becomes high enough, fusion via the triple-alpha process rate starts throughout the core. The core is unable to expand in response to the increased energy production until the pressure is high enough to lift the degeneracy. As a consequence, the temperature increases, causing an increased reaction rate in a positive feedback cycle that becomes a runaway reaction. This process, known as the helium flash, lasts a matter of seconds but burns 60–80% of the helium in the core. During the core flash, the star's energy production can reach approximately 10^{11} solar luminosities which is comparable to the luminosity of a whole galaxy,^[9] although no effects will be immediately observed at the surface, as it is hidden by the star's overlying layers.

For higher mass stars, carbon collects in the core, displacing the helium to a surrounding shell where helium burning occurs. In this helium shell, the pressures are lower and the mass is not supported by electron degeneracy. Thus, as opposed to the center of the star, the shell is able to expand in response to increased thermal pressure in the helium shell. Expansion cools this layer and slows the reaction, causing the star to contract again. This process continues cyclically, and stars undergoing this process will have periodically variable radius and power production. These stars will also lose material from their outer layers as they expand and contract.

Discovery

The triple alpha process is highly dependent on carbon-12 and beryllium-8 having resonances with slightly more energy than helium-4, and before 1952, no such energy levels were known for carbon. The astrophysicist Fred Hoyle used the fact that carbon-12 is abundant in the universe as evidence for the existence of a carbon-12 resonance. The only way Hoyle could find that would produce an abundance of both carbon and oxygen is through a triple alpha process with a carbon-12 resonance near 7.68 MeV.^[10]

Hoyle went to nuclear physicist William Alfred Fowler's lab at Caltech and said that there had to be a resonance of 7.68 MeV in the carbon-12 nucleus. (There had been reports of an excited state at about 7.5 MeV.^[10]) Fred Hoyle's audacity in doing this is remarkable, and initially the nuclear physicists in the lab were skeptical. Finally, a junior physicist, Ward Whaling, fresh from Rice University, who was looking for a project decided to look for the resonance. Fowler gave Whaling permission to use an old Van de Graaff generator that was not being used. Hoyle was back in Cambridge when his prediction was verified a few months later. The nuclear physicists put Hoyle as first author on a paper delivered by Whaling at the Summer meeting of the American Physical Society. A long and fruitful collaboration between Hoyle and Fowler soon followed, with Fowler even coming to Cambridge.^[11] By 1952, Fowler had noted the beryllium-8 resonance, and Edwin Salpeter calculated the reaction rate taking this resonance into account.^{[12][13]}

This helped to explain the rate of the process, but the rate calculated by Salpeter seemed too low at the temperatures expected in supernovas.^[10] When Fowler's lab discovered a carbon-12 resonance near 7.65 MeV it eliminated the discrepancy between the nuclear theory and the theory of stellar evolution.

The final reaction product lies in a 0^+ state (spin 0 and positive parity). Since the Hoyle state was predicted to be either a 0^+ or a 2^+ state, electron-positron pairs or gamma rays were expected to be seen. However, when experiments were carried out, the gamma emission reaction channel was not observed, and this meant the state must be a 0^+ state. This state completely suppresses single gamma emission, since single gamma emission must carry away at least 1 unit of angular momentum. Pair production from an excited 0^+ state is possible because their combined spins (0) can couple to a reaction that has a change in angular momentum of 0.^[14]

Improbability and fine-tuning

Carbon is a necessary component of all known life. ^{12}C , a stable isotope of carbon, is abundantly produced in stars due to three factors:

1. The decay lifetime of a ^8Be nucleus is four orders of magnitude larger than the time for two ^4He nuclei (alpha particles) to scatter.^[15]
2. An excited state of the ^{12}C nucleus exists a little (0.3193 MeV) above the energy level of $^8\text{Be} + ^4\text{He}$. This is necessary because the ground state of ^{12}C is 7.3367 MeV below the energy of $^8\text{Be} + ^4\text{He}$. Therefore, a ^8Be nucleus and a ^4He nucleus cannot reasonably fuse directly into a ground-state ^{12}C nucleus. The excited Hoyle state of ^{12}C is 7.656 MeV above the ground state of ^{12}C . This allows ^8Be and ^4He to use the kinetic energy of their collision to fuse into the excited ^{12}C , which can then transition to its stable ground state. According to one calculation, the energy level of this excited state must be between about 7.3 and 7.9 MeV to produce sufficient carbon for life to exist, and must be further "fine-tuned" to between 7.596 MeV and 7.716 MeV in order to produce the abundant level of ^{12}C observed in nature.^[16]
3. In the reaction $^{12}\text{C} + ^4\text{He} \rightarrow ^{16}\text{O}$, there is an excited state of oxygen which, if it were slightly higher, would provide a resonance and speed up the reaction. In that case, insufficient carbon would exist in nature; almost all of it would have converted to oxygen.^[15]

Some scholars argue the 7.656 MeV Hoyle resonance, in particular, is unlikely to be the product of mere chance. Fred Hoyle argued in 1982 that the Hoyle resonance was evidence of a "superintellect";^[10] Leonard Susskind in The Cosmic Landscape rejects Hoyle's intelligent design argument.^[17] Instead, some scientists believe that different universes, portions of a vast "multiverse", have different fundamental constants:^[18] according to this controversial fine-tuning hypothesis, life can only evolve in the minority of universes where the fundamental constants happen to be fine-tuned to support the existence of life. Other scientists reject the hypothesis of the multiverse on account of the lack of independent evidence.^[19]

References

1. Appenzeller; Harwit; Kippenhahn; Strittmatter; Trimble, eds. (1998). *Astrophysics Library* (3rd ed.). New York: Springer.
2. Carroll, Bradley W. & Ostlie, Dale A. (2007). *An Introduction to Modern Stellar Astrophysics*. Addison Wesley, San Francisco. ISBN 978-0-8053-0348-3.
3. Audi, Georges; Kondev, Filip G.; Wang, Meng; Huang, Wen Jia; Naimi, Sarah (2017), "The NUBASE2016 evaluation of nuclear properties" (<https://www-nds.iaea.org/amdc/ame2016/NUBASE2016.pdf>) (PDF), *Chinese Physics C*, **41** (3): 030001–1–030001–138, Bibcode:2017ChPhC..41c0001A (<http://adsabs.harvard.edu/abs/2017ChPhC..41c0001A>), doi:10.1088/1674-1137/41/3/030001 (<https://doi.org/10.1088%2F1674-1137%2F41%2F3%2F030001>)
4. Wilson, Robert (1997). "Chapter 11: The Stars – their Birth, Life, and Death". *Astronomy through the ages the story of the human attempt to understand the universe*. Basingstoke: Taylor & Francis. ISBN 9780203212738.
5. For example, John Barrow; Frank Tipler (1986). *The Anthropic Cosmological Principle*.
6. Fred Hoyle, "The Universe: Past and Present Reflections." *Engineering and Science*, November, 1981. pp. 8–12
7. Pian, E.; d'Avanzo, P.; Benetti, S.; Branchesi, M.; Brocato, E.; Campana, S.; Cappellaro, E.; Covino, S.; d'Elia, V.; Fynbo, J. P. U.; Getman, F.; Ghirlanda, G.; Ghisellini, G.; Grado, A.; Greco, G.; Hjorth, J.; Kouveliotou, C.; Levan, A.; Limatola, L.; Malesani, D.; Mazzali, P. A.; Melandri, A.; Møller, P.; Nicastro, L.; Palazzi, E.; Piranomonte, S.; Rossi, A.; Salafia, O. S.; Selsing, J.; et al. (2017). "Spectroscopic identification of r-process nucleosynthesis in a double neutron-star merger". *Nature*. **551** (7678): 67–70. arXiv:1710.05858 (<https://arxiv.org/abs/1710.05858>). doi:10.1038/nature24298 (<https://doi.org/10.1038%2Fnature24298>). PMID 29094694 (<https://www.ncbi.nlm.nih.gov/pubmed/29094694>).
8. Carroll, Bradley W.; Ostlie, Dale A. (2006). *An Introduction to Modern Astrophysics* (2nd ed.). Addison-Wesley, San Francisco. pp. 312–313. ISBN 978-0-8053-0402-2.
9. Carroll, Bradley W.; Ostlie, Dale A. (2006). *An Introduction to Modern Astrophysics* (2nd ed.). Addison-Wesley, San Francisco. pp. 461–462. ISBN 978-0-8053-0402-2.
10. Kragh, Helge (2010) When is a prediction anthropic? Fred Hoyle and the 7.65 MeV carbon resonance. <http://philsci-archive.pitt.edu/5332/>
11. Fred Hoyle, *A Life in Science*, Simon Mitton, Cambridge University Press, 2011, pages 205–209.
12. Salpeter, E. E. (1952). "Nuclear Reactions in Stars Without Hydrogen". *The Astrophysical Journal*. **115**: 326–328. Bibcode:1952ApJ...115..326S (<http://adsabs.harvard.edu/abs/1952ApJ...115..326S>). doi:10.1086/145546 (<https://doi.org/10.1086%2F145546>).
13. Salpeter, E. E. (2002). "A Generalist Looks Back". *Annu. Rev. Astron. Astrophys.* **40**: 1–25. Bibcode:2002ARA&A..40....1S (<http://adsabs.harvard.edu/abs/2002ARA&A..40....1S>). doi:10.1146/annurev.astro.40.060401.093901 (<https://doi.org/10.1146%2Fannurev.astro.40.060401.093901>).
14. Cook, CW; Fowler, W.; Lauritsen, C.; Lauritsen, T. (1957). "12B, 12C, and the Red Giants". *Physical Review*. **107** (2): 508–515. Bibcode:1957PhRv..107..508C (<http://adsabs.harvard.edu/abs/1957PhRv..107..508C>). doi:10.1103/PhysRev.107.508 (<https://doi.org/10.1103%2FPhysRev.107.508>).
15. Uzan, Jean-Philippe (April 2003). "The fundamental constants and their variation: observational and theoretical status". *Reviews of Modern Physics*. **75** (2): 403–455. arXiv:hep-ph/0205340 (<https://arxiv.org/abs/hep-ph/0205340>). Bibcode:2003RvMP...75..403U (<http://adsabs.harvard.edu/abs/2003RvMP...75..403U>). doi:10.1103/RevModPhys.75.403 (<https://doi.org/10.1103%2FRevModPhys.75.403>).
16. Livio, M.; Hollowell, D.; Weiss, A.; Truran, J. W. (27 July 1989). "The anthropic significance of the existence of an excited state of 12C". *Nature*. **340** (6231): 281–284. Bibcode:1989Natur.340..281L (<http://adsabs.harvard.edu/abs/1989Natur.340..281L>). doi:10.1038/340281a0 (<https://doi.org/10.1038%2F340281a0>).
17. Peacock, John (2006). "A Universe Tuned for Life". *American Scientist*. **94** (2): 168–170. JSTOR 27858743 (<http://www.jstor.org/stable/27858743>).
18. "Stars burning strangely make life in the multiverse more likely" (<https://www.newscientist.com/article/2104223-stars-burning-strangely-make-life-in-the-multiverse-more-likely/>). *New Scientist*. 1 September 2016. Retrieved 15 January 2017.

19. Barnes, Luke A. "The fine-tuning of the universe for intelligent life." *Publications of the Astronomical Society of Australia* 29.4 (2012): 529–564.

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